Interfacial Oxide Formation Limits the Photovoltage of α -SnWO₄/NiO_x Photoanodes Prepared by Pulsed Laser Deposition

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 α -SnWO₄ is a promising metal oxide photoanode material for direct photoelectrochemical water splitting. With a band gap of 1.9 eV, it ideally matches the requirements as a top absorber in a tandem device theoretically capable of achieving solar-to-hydrogen (STH) efficiencies above 20%. It suffers from photoelectrochemical instability, but NiOx protection layers have been shown to help overcome this limitation. At the same time, however, such protection layers seem to reduce the photovoltage that can be generated at the solid/electrolyte junction. In this study, an extensive analysis of the α -SnWO₄/NiO_x interface is performed by synchrotron-based hard X-ray photoelectron spectroscopy (HAXPES). NiO_x deposition introduces a favorable upwards band bending, but also oxidizes Sn²⁺ to Sn⁴⁺ at the interface. By combining the HAXPES data with open circuit potential (OCP) analysis, density functional theory (DFT) calculations, and Monte Carlo-based photoemission spectra simulation, the presence of a thin oxide layer at the α -SnWO₄/NiO_x interface is suggested and shown to be responsible for the limited photovoltage. Based on this new-found understanding, suitable mitigation strategies can be proposed. Overall, this study demonstrates the complex nature of solid-state interfaces in multi-layer photoelectrodes, which needs to be unraveled to design efficient heterostructured photoelectrodes for solar water splitting.

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DOI: 10.1002/aenm.202003183

1. Introduction

In order to establish a clean and sustainable future energy supply, significant research efforts have been devoted to developing novel renewable energy technologies. Sunlight is the most abundant renewable energy source, but its availability is limited by the diurnal and seasonal cycles. To overcome this, solar energy can be photoelectrochemically converted into hydrogen,^[1,2] which can be stored and used directly as fuel or used as a feedstock to synthesize other chemical fuels.^[3-5] Metal oxide semiconductors have been considered as viable photoelectrode materials for direct photoelectrochemical water splitting.^[6,7] They are generally more stable in aqueous media and easier to process than conventional semiconductors.^[8,9] The semiconducting properties, such as carrier transport, are rather poor (compared to, e.g., III-V^[10-12] semiconductors or Si^[13]), but various strategies have been proposed to at least partially overcome this (e.g., doping, nano-

structuring, overlayer deposition).^[14–16] Using these strategies photocurrents close to the theoretical maximum have been demonstrated for BiVO₄ photoelectrodes, establishing BiVO₄ as one of the highest performing metal oxide photoabsorbers.^[15] This means that the performance of BiVO₄ is no longer limited by its quantum efficiency, but by its relatively large bandgap (\approx 2.4 eV) and the photovoltage losses in the material.

Another metal oxide candidate is α -SnWO₄. There are currently only a few comprehensive studies available on this material, but the reported properties are promising.^[17–24] It is an n-type semiconductor with an indirect band gap of 1.9 eV,^[17] implying a theoretical maximum photocurrent of up to \approx 17 mA cm⁻² under AM 1.5 solar irradiation. Another attractive property of this material is the flat band potential which is located at \approx 0 V versus the reversible hydrogen electrode potential (RHE).^[22–24] To date, α -SnWO₄ thin film photoelectrodes have been prepared successfully using various techniques, including hydrothermal conversion,^[19,23] magnetron sputtering,^[20–22] and pulsed laser deposition (PLD).^[17,18] A major challenge with this material is the limited stability due to surface oxidation of Sn²⁺ to Sn⁴⁺ under (photo)electrochemical



conditions, which presumably results in the formation of an SnO_2 surface layer. This layer blocks charge transport and limits the photocurrent, but it also protects the material from further corrosion (self-passivation). Interestingly, the formation of this blocking layer can be (partially) suppressed by the deposition of a 20 nm NiO_x protection layer using PLD.^[17] This improves the photocurrent, resulting in a record value of 0.75 mA cm⁻² for sulfite oxidation at 1.23 V versus RHE, while still showing good stability.

The improved photocurrent, however, is accompanied by a reduction of the photovoltage, as suggested by the shift in the photocurrent onset potential and the smaller difference in open circuit potential ($\triangle OCP$) between dark and light conditions. We tentatively attributed these limitations to possible Fermi-level pinning at the α -SnWO₄/NiO_x interface, but have not yet been able to show direct evidence for this. In the present study, we provide a detailed fundamental systematic investigation of the α -SnWO₄/NiO_x interface. Synchrotron-based hard X-ray photoelectron spectroscopy (HAXPES) is used to reveal depthdependent information about the chemical states and the electronic structure at the α -SnWO₄/NiO₂ interface. The HAXPES data are complemented with $\triangle OCP$ analysis, density functional theory (DFT) calculations, and Monte Carlo-based photoemission peak intensity simulation. Based on this comprehensive analysis we are able to unravel the origin of the limitation at the α -SnWO₄/ NiO_x interface and propose suitable mitigation strategies.

2. Results and Discussion

We first verify that the NiO_x protection layer limits the photovoltage of α -SnWO₄ by performing a series of Δ OCP measurements. A series of 100 nm-thick phase-pure α -SnWO₄ films (see Figure S1, Supporting Information, for a representative X-ray diffractogram) deposited on FTO were coated with varying thicknesses of NiO_x, using our previously published PLD procedure.^[17,18] **Figure 1** shows the Δ OCP for a pristine α -SnWO₄ film and α -SnWO₄ films coated with increasing thicknesses of NiO_x measured in a 0.5 m KP₁ + 0.5 m Na₂SO₃ electrolyte at pH \approx 72. The Na₂SO₃ served as a hole scavenger. A continuous decrease in Δ OCP is observed with increasing NiO_x thickness up to 10 nm, beyond which the Δ OCP saturates at \approx 0.1 V. This confirms our previous observation^[17] and suggests that the underlying mechanism of the photovoltage limitation is related to a modification of the SnWO₄/NiO_x interface upon deposition of NiO_x.

A reduction of photovoltage in metal oxides is often correlated with the presence of intraband states,^[25,26] which may arise from the formation of hydroxide or oxyhydroxide species at the surface, for instance as a result of air-exposure. Our previously established procedure includes exposing the α -SnWO₄ films to air prior to NiO_x deposition; the films were taken out of the PLD chamber to a separate Ar annealing furnace and reintroduced back into the PLD chamber for NiO_x deposition. To evaluate whether this may be the underlying reason for the photovoltage losses, an alternative preparation process was developed where all deposition and annealing steps are performed within the PLD system (i.e., no air-exposure). Although the dark currents vary, films prepared using both procedures show the same photocurrent and relatively high onset potential (see





Figure 1. Difference of open circuit potential between AM1.5 illumination and dark (Δ OCP) for a pristine α -SnWO₄ and for films coated with different PLD NiO_x layer thicknesses. A decrease of the Δ OCP from 0.22 V (pristine film) to \approx 0.1 V can be observed which saturates after deposition of NiO_x thicker than 10 nm. Measurements were performed in 0.5 m KP_i with 0.5 m Na₂SO₃ added as hole scavenger at pH \approx 7.2. For the NiO_x thickness an error of 10% was estimated from the deposition rate calibration.

Figure S2, Supporting Information, for chopped linear sweep voltammetry), suggesting that air-exposure is not the origin for the photovoltage losses at the α -SnWO₄/NiO_x interface.

A detailed investigation of the chemical states and electronic structure of the α -SnWO₄/NiO_x interface was performed using synchrotron-based HAXPES. Such a study will help to identify the root cause of the interfacial losses and propose appropriate improvement strategies. Our experimental approach is depicted in **Figure 2**.

In brief, similar to the $\triangle OCP$ measurements, we varied the NiO_x layer thickness (between 0 and 50 nm) deposited on the 100 nm α -SnWO₄ film. As the NiO_x thickness increases, the photoelectron signal originating from α -SnWO₄ is increasingly attenuated until only photoelectrons escaping from the NiO_x layer are detected. This allows us to de-convolute the interfacial spectral features from those arising from the bulk and/or the surface. In addition, depth-dependent information is further gained by varying the kinetic energy of the incident photons (hv) between 2, 4, and 6 keV. The respective probing depths, σ , for these different photon energies correspond to three times the inelastic mean free path, λ , and are indicated in Figure 2.

Before we discuss the effect of NiO_x deposition in detail, the chemical states of pristine α -SnWO₄ films are first evaluated. W 4f and Sn 3d core level spectra of a pristine α -SnWO₄ film measured with $h\nu = 2$ keV are shown in **Figure 3**a,b, respectively. The W 4f core level spectrum can be deconvoluted into two contributions (Figure 3a, see Experimental Section for the fitting procedure). W⁶⁺ is mainly present as expected for α -SnWO₄, and a minor contribution of W⁵⁺ is also detected. We confirm that the existence of W⁵⁺ is not related to beam damage (as is known for WO₃^[27]), as shown from our measurements with different beam filters and at different sample spots (see Figure S3, Supporting Information, for details). The contribution of W⁵⁺ to the





NiO_x layer thickness (2 – 50 nm)

Figure 2. Schematic illustration of our HAXPES interface experiment. α -SnWO₄ films are coated with a NiO_x layer with increasing thickness such that the interfacial spectral features can be deconvoluted from features arising from the bulk and the surface. The kinetic energy of photons (h ν) is varied between 2, 4, and 6 keV. The increasing probing depth σ , which is equal to three times the inelastic mean free path λ , for the different photon energies is indicated. The brown arrows depict the excited photoelectrons from α -SnWO₄, while the excited photoelectrons from NiO_x are illustrated with the clay creek arrows.

total signal increases with increasing photon energy hv, which corresponds to increasing probing depth σ (see Figure 3c; the spectra are shown in Figure S4a, Supporting Information). This suggests that W⁵⁺ is present more in the bulk than at the surface of the film. We note that W5+ was not observed in the lab-based X-ray photoelectron spectroscopy (XPS) spectra of our previously reported α -SnWO₄,^[17] whereas the same lab-based XPS measurement for our current sample does reveal the presence of W⁵⁺. Since samples produced both in the previous and current work have very similar $\triangle OCP$ and photocurrent onset potential, we confirm that the presence of W^{5+} has no impact on these parameters, which are the metrics of interest investigated here. The exact origin of the W⁵⁺ contribution observed in the present study is not fully clear at this point and while the presence of W5+ is undoubtedly interesting-for example, to investigate the influence on the charge transport properties-it is not the main motivation of the present study. The Sn 3d core level spectrum (Figure 3b) shows the presence of mainly Sn²⁺ with a small contribution of oxidized Sn^{4+} . For hv = 2 keV, the relative contribution of Sn4+ to the total signal is estimated to be $18 \pm 3\%$, which is in the same range as measured previously with lab-based photoelectron spectroscopy (hv = 1486.74 eV).^[17] In contrast to W⁵⁺, the relative contribution of Sn⁴⁺ tends to decrease with increasing probing depths σ (Figure 3c; see Figure S4b, Supporting Information, for the spectra). This indicates that the relative contribution of Sn4+ is higher at the surface of α -SnWO₄ than in the bulk of the material, which makes it relevant for our interface study. Finally, we note that the binding energies do not shift with varying photon energies (see Figure S4, Supporting Information), indicating that no significant band bending is present at the surface of pristine α -SnWO₄ films under the HAXPES measurement conditions.

The influence of the NiO_x film on the α -SnWO₄/NiO_x interface is now evaluated by considering the evolution of the W 4f and Sn 3d core level spectra with increasing NiO_x thickness. **Figure 4**a shows the W 4f core level spectra for SnWO₄ films coated with different NiO_x layer thicknesses, measured with $h\nu = 2$ keV (see Figures S6–S8, Supporting Information, for fitted spectra, also at different photon energies). The W⁶⁺ peaks show a clear shift to lower binding energies with increasing NiO_x thickness (see vertical red dashed line in Figure 4a). A detailed discussion on this will follow later. The contribution of W⁵⁺ is still clearly present for films with NiO_x thickness up to 10 nm and the position of the peak (i.e., binding energy) remains relatively constant independent of the photon energy (see Figure 4c and the vertical orange dashed line at 33.8 eV in Figure 4a). For thicknesses larger than 10 nm the relative W^{5+} contribution seems to decrease, although the poor signal-to-noise ratio precludes a quantitative evaluation. This trend is especially true for the data obtained using higher photon energies (i.e., 4 and 6 keV, see Figure 4c), which agrees very well with the notion that W^{5+} is present more in the bulk versus at the surface/interface.

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In the following, the influence of the NiO_x thickness on the Sn 3d core level spectra (Figure 4b) is evaluated. All spectra can be de-convoluted into Sn²⁺ and Sn⁴⁺ contributions (see Figures S9-S11, Supporting Information). The relative Sn⁴⁺ contribution increases with increasing NiOx layer thickness (Figure 4d), and for a photon energy of 2 keV, Sn⁴⁺ becomes the dominant oxidation state (>50%) when the NiO_r thickness is ≥ 10 nm. The increasing contribution of Sn⁴⁺ is also apparent from the shifts of the peak maxima towards higher binding energies (Figure 4b). We note that this trend qualitatively correlates with the trend of $\triangle OCP$ shown in Figure 1, that is, the $\triangle OCP$ starts to saturate at the same NiO_x layer thickness when Sn4+ starts to dominate the HAXPES signal. The same trend is observed for the different photon energies (Figure 4d), but the higher photon energies (i.e., higher probing depths) show a smaller increase of the relative Sn⁴⁺ contribution with increasing NiO_x thickness. This indicates that the oxidation of Sn occurs at the α -SnWO₄/NiO_x interface and not in the bulk of α -SnWO₄. The influence of the increasing contribution of Sn⁴⁺ on the interface properties and the photovoltage will be discussed in more detail below.

The O1s core level spectra measured with different photon energies can be found in Figures S12–S14, Supporting Information. The ratios between peaks arising from oxygen in α -SnWO₄ and NiO_x (at higher and lower binding energies, respectively) change in the expected manner with the NiO_x thickness and the photon energy hv. The Ni 2p core level spectra do not change significantly for different NiO_x layer thicknesses, independent of the photon energy (see Figures S12-S15, Supporting Information). This indicates that the chemical nature of NiO_x remains the same with increasing thickness. The valence band (VB) spectra were recorded as well and can be seen in Figure S16,







Figure 3. a) W 4f and b) Sn 3d core level photoemission spectra for pristine α -SnWO₄ films measured with $h\nu = 2$ keV. A contribution of reduced W⁵⁺ can be observed in W 4f core levels in addition to W⁶⁺ corresponding to nominal α -SnWO₄. In Sn 3d core levels a small contribution of oxidized Sn⁴⁺ differing from nominal Sn²⁺ can be found as indicated by the peak asymmetry towards higher binding energy. c) Relative W⁵⁺ and Sn⁴⁺ contributions to the total W 4f and Sn 3d signals, respectively, as a function of photon energy, extracted by peak fitting (the spectra are shown in Figure S4, Supporting Information). The probing depth σ corresponds to three times the inelastic mean free path λ at each particular kinetic energy.

Supporting Information, with features of the α -SnWO₄ VB becoming smaller, as expected, with increasing NiO_x thickness.

We note that our ability to still detect photoelectrons from the W 4f and Sn 3d core levels for the thicker NiO_x films (>10 nm) seems inconsistent with the expected probing depth of the photon energies used. This is because the NiO_x films are not flat and dense, but instead exhibit island-type growth, as already shown by the scanning electron micrographs in our earlier publication.^[17]

The binding energies for different peaks extracted from fitting the W $4f_{7/2}$ and Sn $3d_{5/2}$ core level spectra are shown in **Figure 5** as a function of the NiO_x layer thickness (see

Figures S6–S11, Supporting Information, for fitted spectra). The data for the different oxidation states (i.e., W^{5+} , W^{6+} , Sn^{2+} , Sn^{4+}) are shown separately. A shift of ≈ 0.4 eV towards lower binding energies is observed for W^{6+} . In contrast, as already mentioned above, NiO_x deposition has no significant impact on the binding energies of W^{5+} . The Sn $3d_{5/2}$ peaks also experience a shift towards lower binding energies and saturate after the deposition of 10 nm NiO_x. Both Sn²⁺ and Sn⁴⁺ peaks experience the same ≈ 0.4 eV shift that is observed for the W⁶⁺ peak.

We attribute the correlated shift in the binding energies of W^{6+} , Sn^{2+} , and Sn^{4+} to an electronic shift and a change in band bending in α -SnWO₄ as a result of NiO_x deposition. A shift to



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Figure 4. a) W 4f and b) Sn 3d core level spectra measured with hv = 2 keV for a pristine α -SnWO₄ film (black) and films coated with increasing NiO_x layer thickness (blue). The dashed vertical lines indicate the binding energies of the W⁵⁺ (red), W⁶⁺ (orange), and Sn²⁺ (dark cyan) of the pristine α -SnWO₄ film. O 1s, Ni 2p, and valence band spectra can be found in Figures S12–S16, Supporting Information, and the W 4f and Sn 3d spectra measured with higher h v are shown in Figure S5, Supporting Information. c) The relative W⁵⁺ contribution to the total signal in dependence on the NiO_x layer thickness seems to decrease, especially based on the data obtained using 4 and 6 keV photon energies (see inset), although the poor signal-to-noise ratio prevents a quantitative analysis for samples with thicker NiO_x. The corresponding data points for samples with 20 and 50 nm NiO_x can therefore not be depicted. d) The relative contribution of Sn⁴⁺ increases with increasing NiO_x layer thickness. The increase of the relative Sn⁴⁺ contribution is larger for lower h v indicating that the oxidation occurs at the interface and not the bulk. For the sample with 50 nm NiO_x a quantitative analysis is not possible for a photon energy of 6 keV due to the poor signal-to-noise ratio.

lower binding energies corresponds to an upward bending of the energy bands (i.e., additional band bending at the surface). At the same time, the Ni 2p binding energies do not shift (see Figures S12-S15, Supporting Information), indicating that the additional band bending is only present in the α -SnWO₄ film. Similar observations have been reported in the literature, for instance for BiVO₄ deposited with various overlavers^[28,29] and metal oxide/Si buried junction photoabsorbers,^[30,31] where the observed performance improvement has been attributed to additional band bending. In our case, however, the additional band bending with increasing NiO_x deposition contradicts the decreasing \triangle OCP shown in Figure 1. This suggests that a compensating factor is present that negates the favorable effect of band bending. We assign the significant oxidation of Sn²⁺ to Sn⁴⁺ as the compensating factor that causes the reduction of the photovoltage. The details of this mechanism are discussed in the following paragraphs.

We should point out that the lack of binding energy shift for the W⁵⁺ peaks seems inconsistent with the band bending explanation above. While the reason for this is not clear at this point, we speculate that it may be caused by one (or a combination) of the following two factors. First, the W⁵⁺ species may be present in a separate amorphous impurity phase outside of the α -SnWO₄ phase. Second, since band bending is a pure interface phenomenon, the peak position of W⁵⁺—which is predominantly present in the bulk—is less sensitive to band bending than other species at the interface. We note, however, that the presence of W⁵⁺ does not affect the limited photovoltage in α -SnWO₄ which is the focus of this study.

DFT/HSE06 method calculations were applied to analyze the electronic structure of α -SnWO₄ and how this is influenced by the presence of Sn⁴⁺ defects. We used a similar methodology as has been reported previously for pristine α -SnWO₄.^[21,22] We took a compensation approach by introducing vacancies in





Figure 5. Influence of NiO_x layer thickness on a) W $4f_{7/2}$ and b) Sn $3d_{5/2}$ peak positions measured with different photon energies h ν . No shift is observed for W⁵⁺, while the W⁶⁺, Sn²⁺, and Sn⁴⁺ peaks shift to lower binding energies (by up to ≈ 0.4 eV) with increasing NiO_x layer thickness. This is attributed to upwards band bending in α -SnWO₄. All binding energies were determined by the fitting procedure described in the Experimental Section.

order to simulate the presence of Sn⁴⁺. Among the explored selfdefective structures containing either O-, Sn- or W-vacancies and the structures containing anti-site defects, only the one associated with W-vacancy could represent the possible presence of Sn⁴⁺ (see Figures S17 and S18, Supporting Information, for more details regarding other structures). Figure 6a,b show the orthorhombic crystal structure of pristine α -SnWO₄ as well as α -SnWO₄ that contains Sn⁴⁺ defects. The density of states (DOS) plots for these two structures are shown in Figures 6c and 6d. The band gap in α -SnWO₄ originates from indirect transitions between the occupied Sn $5s^2 + O 2p^6$ and the unoccupied W 5d⁰ states. The introduction of Sn⁴⁺ does not modify the band gap, but it results in the appearance of a narrow state in the middle of the band gap. This state has a highly localized character as revealed by the partial charge density map shown in Figure S19, Supporting Information. Such a mid-gap state is known to be able to pin the Fermi level and therefore restrict the extractable photovoltage.^[32,33]

Before making any definite conclusions based on these DFT results, we first critically assess how well the structure in the proposed model represents our α -SnWO₄ films. In particular, it has to be examined whether W vacancies, required to compensate the Sn⁴⁺ defects, actually occur in our α -SnWO₄ films. There are three arguments against the presence of W vacancies. First, we observed previously by Rutherford backscattering spectrometry (RBS) that our deposition procedure generally results in slightly W-rich films.^[17] Second, based on the compensation model used in the DFT, one would expect a higher value of the Sn/W ratio with increasing NiO_x thickness, that is, more W vacancies due to a higher amount of Sn⁴⁺. However, this is inconsistent with our HAXPES data (see Figure S20, Supporting Information). Finally, for electrostatic reasons, a W vacancy is likely to have a high defect formation enthalpy due to the high formal oxidation state of W (6+). All these arguments suggest that it is rather unlikely that a significant amount of W vacancies is present in our samples. Consequently, we conclude that there has to be another mechanism that chargecompensates the Sn⁴⁺ defects and that may explain the reduced photovoltage.

Another mechanism that may explain the increase of Sn⁴⁺ is the formation of a secondary phase at the α -SnWO₄/NiO_x interface. Although the NiO_x deposition was performed at room temperature, the presence of O₂ plasma during the PLD process may cause the formation of this phase; indeed, a preliminary experiment of intentionally exposing our pristine α -SnWO₄ to O₂ plasma resulted in a similar increase of Sn⁴⁺ contribution. Such a formation of interfacial layer from a secondary phase has been reported in other systems, such as at the Cu₂O/ZTO (zinc-tin-oxide) interface and Si/HfO2 interface.^[34-36] A possible identity of this secondary phase is Sn0.23WO3, which has been reported to be present as impurities in several *a*-SnWO₄ films.^[18,37] This is, however, unlikely to be the case. The Sn_{0.23}WO₃ phase has been reported to contain Sn²⁺ and Sn⁰, but no Sn^{4+} .^[38] $Sn(WO_4)_2$ is another possibility, but to the best of our knowledge, this phase has only been reported based on theoretical calculation.^[39]

Instead, we speculate that the secondary phase at the α -SnWO₄/NiO_x interface is SnO₂. We thus quantitatively assess whether the formation of an SnO₂ layer can be correlated with our HAXPES results by performing Monte Carlo simulations of photoemission peak intensities using the SESSA software package.^[40] The simulations were performed for α -SnWO₄ films coated with an increasing thickness of SnO₂ at the surface (see Figure S21, Supporting Information, for more details). For simplicity, we do not include the NiO_x layer in the simulations, since it would only attenuate the overall Sn 3d spectrum but not change the relative contribution of Sn²⁺ and Sn⁴⁺ to the spectra, regardless of where these species are located (surface or bulk).

Figure 7a displays the relative contribution of Sn^{4+} to the total Sn 3d HAXPES signal as a function of NiO_x thickness, as obtained from our experimental data. This data is now compared with the simulated results. For each simulated spectrum of an α -SnWO₄ film covered with different SnO₂ layer thicknesses (see Figure S21, Supporting Information), the relative contribution of Sn⁴⁺ is extracted and shown as a function of the SnO₂ layer thickness in Figure 7b. We combine this simulated data with the experimentally observed relative Sn⁴⁺ contribution





Figure 6. DFT-based orthorhombic crystal lattice of a) pristine α -SnWO₄ and b) α -SnWO₄ containing W vacancies to compensate Sn⁴⁺ defects. The corresponding density of states (DOS) is shown in (c) and (d), respectively. In the case where W vacancies are present, the triplet spin state is the most stable configuration; the spin-up (dashed lines) and spin-down (solid lines) components of the DOS are therefore shown.

(Figure 7a) in order to estimate the expected interfacial SnO_2 layer thickness for films with various deposited NiO_x layer thickness. The resulting data is shown in Figure 7c, which clearly shows that a thicker NiO_x layer indeed corresponds to a thicker SnO_2 layer at the interface.

For the sample that shows record photocurrent (i.e., 20 nm thick NiO_x), an interfacial SnO₂ layer thickness of \approx 2 nm is estimated. The transmission electron microscope (TEM) image presented in Figure S22, Supporting Information, suggests that such a layer is indeed present at the interface. Although the interfacial layer is thin enough to allow charge carriers to be transferred across the interface (i.e., tunneling), such a layer may negatively affect the α -SnWO₄/NiO_x interface. For example, this insulating layer may limit the extractable photocurrent. An important question relevant to the current study is whether the presence of such a layer can be correlated with the reduced photovoltage. To address this question, Δ OCP measurements were performed again with a pristine α -SnWO₄ film,

to which different durations of chronoamperometry treatment under AM1.5 illumination at a potential of 1.23 V versus RHE were applied. Since such a treatment has been shown to induce SnO_2 layer formation at the surface of α -SnWO₄,^[17] the different durations (i.e., different amounts of transferred charges across the interface) would result in different thicknesses of the SnO₂ layer. After each step an OCP measurement was performed. **Figure 8** shows the $\triangle OCP$ as a function of the amount of charge transferred across the α -SnWO₄ surface (Q), calculated by integrating the chronoamperometry data. The SnO₂ layer thickness on the surface can also be estimated from the amount of charge and the reported density of SnO_2 (7.02 g cm⁻³),^[41] as shown at the top x-axis of the figure (see also Figure S23, Supporting Information, for the cross-section SEM image for the sample with the thickest SnO_2 layer). With increasing Q, the $\triangle OCP$ decreases reaching a value of ≈ 0.08 V. This confirms that the formation of an SnO₂ layer on the surface of α -SnWO₄ results in a reduction of the \triangle OCP. Based on this observation and the







Figure 7. a) Relative Sn^{4+} contribution to the total signal as a function of NiO_x thickness from HAXPES experimental data measured with a photon energy of 2 keV. b) Relative Sn^{4+} contribution to the total signal as a function of SnO_2 layer thickness extracted from Monte Carlo simulation of peak intensities with the SESSA software package, also simulated with a photon energy of 2 keV. c) Correlation of simulation and experimental data. The SnO_2 layer thickness is estimated by using the fraction of Sn^{4+} signal observed experimentally for a given NiO_x thickness and reading the SnO_2 thickness from (a) for that fraction of Sn^{4+} signal (b). For the sample that shows the record photocurrent (i.e., 20 nm thick NiO_x), an interfacial SnO_2 layer thickness of ≈ 2 nm is estimated.

qualitative agreement with data shown in Figure 1, we conclude that the formation of an interfacial SnO₂ layer at the α -SnWO₄/NiO_x interface is responsible for the limited photovoltage that can be extracted.

We finally note that our current findings complement our previous report,^[17] where it was mentioned that the holeinduced electrochemical oxidation of Sn^{2+} to Sn^{4+} is prevented by a NiO_x protection layer and a stable photocurrent for at least 0.5 h can be achieved. Our more thorough HAXPES investigation in the present study provides additional understanding on



Figure 8. \triangle OCP measurements performed on pristine α -SnWO₄ after being photoelectrochemically treated at 1.23 V versus RHE for a specific duration. The charge Q was calculated from the recorded chronoamperometry data. The decrease in \triangle OCP is correlated with the oxidation of Sn²⁺ to Sn⁴⁺ and the formation of an SnO₂ layer. The estimated SnO₂ layer thickness on the top x-axis was calculated from the transferred charge Qand the reported density of SnO₂ (7.02 g cm⁻³).^[41] The dashed line is a guide for the eye.

the protection mechanism. PLD NiO_x introduces some oxidation of α -SnWO₄ during deposition, but it limits the formed SnO₂ layer thickness such that charge carrier transport is still possible. Furthermore, NiO_x impedes further (photo)electrochemical oxidation of SnWO₄, and the formation of a thick hole blocking SnO₂ layer (>7 nm).

3. Conclusion

In summary, we performed a detailed investigation of the α -SnWO₄/NiO₂ interface using synchrotron-based HAXPES in order to understand the origin of the limited photovoltage in NiO_x-coated α -SnWO₄ films. NiO_x deposition was found to introduce a favorable upwards band bending at the *α*-SnWO₄/ NiO_x interface. However, this is accompanied by the oxidation of Sn²⁺ to Sn⁴⁺, the degree of which is in a good agreement with the decrease of photovoltage as shown by the open circuit potential (OCP) analysis. Using DFT and Monte Carlo based photoemission intensity simulation, we showed that the increase of Sn⁴⁺ signal can be explained by the formation of a thin oxide layer, such as SnO_2 , at the interface of α -SnWO₄ and NiO_x. For the optimized sample with 20 nm-thick NiO_x, an interfacial oxide layer thickness of ≈ 2 nm is estimated. Although this is in principle thin enough for charge carriers to tunnel through, control experiments by intentionally converting the surface of *a*-SnWO₄ films to an SnO₂ layer with different thicknesses indeed revealed a reduction of the photovoltage in this thickness range. This suggests that the Fermi level pinning is not caused by a direct contact of *α*-SnWO₄ and NiO_x, but is instead related to the presence of the interfacial oxide layer. To overcome the photovoltage limitation, alternative deposition techniques for NiO_x (e.g., vacuum evaporation, atomic layer deposition) or other co-catalyst/protection layers that do not alter the surface of oxidation-sensitive α -SnWO₄ need to be explored. This would prevent the formation of the interfacial oxide layer and would enable the extraction of the

maximum photovoltage from the system. Finally, our study shows that understanding the interface properties of metal oxide-based solid-state junctions is paramount in the development of highly efficient and stable electrodes for photoelectrochemical water splitting.

4. Experimental Section

Synthesis of α -SnWO₄ and NiO_x: PLD was used to deposit α -SnWO₄ thin films and NiO_x overlayers based on a previously published procedure,^[17] and is briefly described below. Fluorine-doped tin oxide coated glass (FTO TEC 7, Pilkington) substrates were used as the transparent conductive back contact. Prior to deposition, the substrates (24 × 24 mm²) were successively cleaned in an ultrasonic bath using a 1 vol% Triton solution (Triton X-100, laboratory grade, Sigma-Aldrich), acetone, isopropanol, ethanol, and deionized water for 15 min each. Afterwards they were dried under an N₂ stream.

Ablation of a ceramic α -SnWO₄ target was performed in a custom-built PLD system from PREVAC at room temperature equipped with a KrFexcimer laser with a wavelength of $\lambda = 248$ nm (LPXpro 210, Coherent). The base pressure of the deposition chamber was $\approx\!\!1\times10^{-7}$ mbar. The laser fluence was adjusted to 2 | cm⁻² with a laser spot size of $1.3 \times 2 \text{ mm}^2$ and the repetition rate was 10 Hz. Film deposition was performed with an oxygen background pressure of $\approx 10^{-4}$ mbar. Substrates were fixed in an off-axis position with a target-to-substrate distance of 60 mm. A continuous substrate rotation of 12° s⁻¹ was applied to achieve homogeneous films. Prior to every deposition a pre-ablation routine was performed to transform the target to a steady state condition, which was achieved when a constant deposition rate was monitored with a quartz crystal microbalance (QCM). Different film thicknesses were achieved by adjusting the respective number of laser shots. After deposition, the films were post-annealed in a tube furnace in Ar atmosphere at 520 °C for 2 h to prevent film oxidation. The presence of oxygen was minimized by evacuating the tube and flushing with Ar twice before annealing.

 $\rm NiO_x$ was deposited by ablating a metallic Ni target in the same PLD system. The process was done at room temperature under an oxygen background pressure of 0.05 mbar. The laser fluence was 2 J cm⁻² and a repetition rate of 10 Hz was used.

Characterization: X-ray diffraction was performed with a Bruker D8 Advance diffractometer in a grazing incidence configuration with a grazing incidence angle of 0.5°. The step size was set to 0.02° and the respective integration time per step was 6 s. Cu K α (λ = 1.5406 Å) was used as the X-ray source with a voltage of 40 kV and a current of 40 mA.

Photoelectrochemistry (PEC) measurements were performed in a three-electrode configuration in a custom-made PEC cell. The sample was connected as the working electrode to an EG&G Princeton Applied Research (Model 273A) potentiostat. A platinum wire was used as the counter electrode, and an Ag/AgCl electrode (XR300, saturated KCl, Radiometer Analytical) was used as the reference electrode. Measured potentials were translated to the RHE by using the Nernst equation. AM1.5 illumination was obtained with a WACOM Super Solar Simulator (Model WXS-505-5H, AM1.5, Class AAA). The used electrolyte was 0.5 M potassium phosphate buffer (KPI, pH 7.2) made from dissolving KH_2PO4 (\geq 99.0%, Sigma-Aldrich) and K_2HPO4 (\geq 99.0%, Sigma-Aldrich) with a 1:1.4 molar ratio in Milli-Q water. As a hole scavenger, 0.5 M Na_2SO₃ was added to the buffer solution. OCP measurements were performed under the same conditions.

HAXPES was performed at the Bessy II synchrotron facility at Helmholtz-Zentrum Berlin, Germany. Measurements were performed at the high kinetic energy photoemission end station $(HIKE)^{[42]}$ at the KMC-1 beamline.^[43] During the measurements the analysis chamber was kept at $\approx 10^{-8}$ mbar. Photon energies were varied between 2003, 4000, and 6009 eV using the Si (111) and Si (311) cut crystals pairs of a double crystal monochromator. Photoelectrons were detected by a Scienta R4000 analyzer. Survey spectra were recorded with a step size of 0.5 eV and a pass energy of 200 eV. Core level and valence band spectra were

recorded with a step size of 0.05 eV and a pass energy of 200 eV. Binding energies were referenced to the Au 4f core level (binding energy set equal to 84.00 eV) measured on a gold foil located on the sample manipulator which shared a common ground with the sample and the analyzer; in order to extract the binding energies and the contribution of the respective oxidation states from measured spectra a fitting procedure was applied. Before that a Shirley baseline subtraction was performed. The peak shape was represented by pseudo-Voigt functions consisting of a convolution of Gaussian and Lorentzian functions. Finally, chi-square minimization was performed to optimize the fitted spectra.

TEM was performed using a Philips CM12/STEM. The system is equipped with a LaB₆ cathode and a super twin lens for high resolution. An acceleration voltage of 120 keV was used. Four preparation samples were cut into small pieces and polished from both sides until reaching a thickness of 4–6 μ m. Afterwards a Mo ring was glued on the sample. The samples were further thinned with an Ar ion mill until they had electron transparent areas.

Density Functional Theory: Starting from the $2 \times 1 \times 2$ supercell model, which includes 16 functional units of $Sn_{16}W_{16}O_{64}$ or 96 atoms of the orthorhombic crystal lattice of α -SnWO₄ (*Pnna* space group),^[18,19,22-24] a perfect lattice and five representative self-defective lattices containing intrinsic O-, Sn-, and W-vacancies as well as Sn-antisite on W and W-antisite on Sn were modeled to mimic the possible presence of Sn⁴⁺ in α -SnWO₄ samples. For O-, Sn-, and W-deficient α -SnWO₄, one O atom, one Sn-atom, and one W atom were removed from the Sn₁₆W₁₆O₆₄ supercell, respectively. For Sn-antisite on W, one W-atom was replaced by one Sn-atom in the Sn₁₆W₁₆O₆₄ supercell. For W-antisite on Sn, one Sn-atom was replaced by one W-atom in the Sn₁₆W₁₆O₆₄ supercell. For each defect type, different spin multiplicities were explored in order to come up with the most stable spin configuration.

The various generated crystalline structures were fully optimized by means of the spin-polarized DFT using the Vienna Ab initio simulation package (VASP)^[44-46] with the Perdew–Burke–Ernzerhof (PBE) exchange correlation potential^[47] and the frozen-core projector augmented-wave (PAW) approach.^[48] The configurations of valence electrons treated explicitly in the plane wave descriptions are 5s²5p² for Sn, 5d⁴6s² for W, and 2s²2p⁴ for O. Upon several benchmark tests evaluating the accuracy of the calculations, a kinetic cutoff energy of 400 eV was consistently used for electron wave functions expansion and a Monkhorst-Pack k-point mesh^{[49]} of $3 \times 3 \times 3$ was used for sampling the Brillion zone. Gaussian smearing was used with a sigma value of 0.1 eV. The atomic positions were fully relaxed with the conjugate gradient procedure until the residual forces vanished within 0.01 eV Å-1 and the electronic convergence for each supercell was below 10⁻⁶ eV. Our PBE-based computed lattice parameters (a = 5.59 Å, b = 11.63 Å, c = 4.98 Å, and $\alpha = \beta = \gamma = 90^{\circ}$) of the pristine material are in excellent agreement with the experimental values (a = 5.62 Å, b = 11.74 Å, c = 4.98 Å, and $\alpha = \beta = \gamma = 90^{\circ}$).^[17,22-24,50]

By considering the optimized geometries obtained with the PBE functional, the electronic structure calculations of pristine and the most relevant self-defective α -SnWO₄ materials were conducted at the level of the more expensive screened coulomb hybrid Heyd-Scuseria-Ernzerhof (HSE06)^[51] functional using VASP. This method has been proven to provide much more accurate band gap values of a large number of photocatalytic semiconductors than the standard generalized-gradient approximation (GGA) functional.^[52-58] The electronic properties analysis covered the DOS together with the partial charge density maps corresponding to the new states introduced by the created defects in terms of orbital contribution types of each element in the cell. The validity of our electronic structure calculation is demonstrated by the relatively good agreement between our HSE06-based computed band gap of 1.5 eV for the pristine material with the reported experimental values (in between 1.6 and 1.9 eV), which is much more accurate than the PBE computed one of 0.8 eV.[17,19,22-24]

Supporting Information

Supporting Information is available from the Wiley Online Library or from the author.



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Acknowledgements

The authors acknowledge financial support for this work from the Helmholtz International Research School "Hybrid Integrated Systems for Conversion of Solar Energy" (HI-SCORE), an initiative co-funded by the Initiative and Networking Fund of the Helmholtz Association. Part of the work was funded by the Volkswagen Foundation. They would also like to thank Roberto Felix Duarte and Regan Wilks for the access and technical assistance to the HiKE endstation, KMC-1 beamline at the BESSY-II synchrotron facility, as well as Karsten Harbauer and Ronen Gottesman for supporting the PLD experiments. M.H. thanks the KAUST Supercomputing Laboratory for the needed computational resources. The authors thank Ulrike Bloeck for recording the TEM images.

Open access funding enabled and organized by Projekt DEAL.

Conflict of Interest

The authors declare no conflict of interest.

Data Availability

The data that support the findings of this study are available from the corresponding author upon reasonable request.

Keywords

DFT, HAXPES, interface, photovoltage, pulsed laser deposition, $\alpha\text{-}\mathsf{SnWO}_4$

Received: October 6, 2020 Revised: December 27, 2020 Published online: January 14, 2021

- M. G. Walter, E. L. Warren, J. R. McKone, S. W. Boettcher, Q. Mi, E. A. Santori, N. S. Lewis, *Chem. Rev.* 2010, *110*, 6446.
- [2] M. Grätzel, Nature 2001, 414, 338.
- [3] J. A. Turner, Science 2004, 305, 972.
- [4] R. van de Krol, B. A. Parkinson, MRS Energy Sustain. 2017, 4, E13.
- [5] S. Ardo, D. F. Rivas, M. A. Modestino, V. Schulze Greiving, F. F. Abdi, E. Alcaron Llado, V. Artero, K. Ayers, C. Battaglia, J.-P. Becker, D. Bederak, A. Berger, F. Buda, E. Chinello, B. Dam, V. Di Palma, T. Edvinsson, K. Fujii, H. Gardeniers, H. Geerlings, S. M. H. Hashemi, S. Haussener, F. Houle, J. Huskens, G. F. Morre, J. Muller, K. L. Orchard, T. E. Rosser, F. H. Saadi, J.-W. Schüttauf, B. Seger, S. W. Sheehan, W. A. Smith, J. Spurgeon, M. H. Tang, R. van de Krol, P. C. K. Vesborg, P. Westerik, *Energy Environ. Sci.* 2018. 11. 2768.
- [6] D. K. Lee, D. Lee, M. A. Lumley, K.-S. Choi, Chem. Soc. Rev. 2019, 48, 2126.
- [7] F. F. Abdi, S. P. Berglund, J. Phys. D: Appl. Phys. 2017, 50, 193002.
- [8] K. Sivula, R. van de Krol, Nat. Rev. Mater. 2016, 1, 15010.
- [9] K. Sivula, J. Phys. Chem. Lett. 2013, 4, 1624.
- [10] O. Khaselev, J. A. Turner, Science 1998, 280, 425.
- [11] J. L. Young, M. A. Steiner, H. Döscher, R. M. France, J. A. Turner, T. G. Deutsch, *Nat. Energy* **2017**, *2*, 17028.
- [12] M. M. May, H.-J. Lewerenz, D. Lackner, F. Dimroth, T. Hannappel, *Nat. Commun.* 2015, 6, 8286.
- [13] M. J. Kenney, M. Gong, Y. Li, J. Z. Wu, J. Feng, M. Lanza, H. Dai, Science 2013, 342, 836.



www.advenergymat.de

- [14] T. W. Kim, K.-S. Choi, Science 2014, 343, 990.
- [15] Y. Pihosh, I. Turkevych, K. Mawatari, J. Uemura, Y. Kazoe, S. Kosar, K. Makita, T. Sugaya, T. Matsui, D. Fujita, M. Tosa, M. Kondo, T. Kitamori, *Sci. Rep.* 2015, *5*, 11141.
- [16] A. Kudo, K. Omori, H. Kato, J. Am. Chem. Soc. 1999, 121, 11459.
- [17] M. Kölbach, I. J. Pereira, K. Harbauer, P. Plate, K. Höflich, S. P. Berglund, D. Friedrich, R. van de Krol, F. F. Abdi, *Chem. Mater.* 2018, 30, 8322.
- [18] M. Kölbach, H. Hempel, K. Harbauer, M. Schleuning, A. Petsiuk, K. Höflich, V. Deinhart, D. Friedrich, R. Eichberger, F. F. Abdi, R. van de Krol, ACS Appl. Energy Mater. 2020, 3, 4320.
- [19] K. J. Pyper, T. C. Evans, B. M. Bartlett, Chin. Chem. Lett. 2015, 26, 474.
- [20] F. Bozheyev, E. M. Akinoglu, L. Wu, H. Lu, R. Nemkayeva, Y. Xue, M. Jin, M. Giersig, *Int. J. Hydrog. Energy* **2020**, *45*, 8676.
- [21] M. Harb, A. Ziani, K. Takanabe, Phys. Status Solidi B 2016, 253, 1115.
- [22] A. Ziani, M. Harb, D. Noureldine, K. Takanabe, APL Mater. 2015, 3, 096101.
- [23] Z. Zhu, P. Sarker, C. Zhao, L. Zhou, R. L. Grimm, M. N. Huda, P. M. Rao, ACS Appl. Mater. Interfaces 2017, 9, 1459.
- [24] I. Cho, C. H. Kwak, D. W. Kim, S. Lee, K. S. Hong, J. Phys. Chem. C 2009, 113, 10647.
- [25] B. Klahr, T. Hamann, J. Phys. Chem. C 2014, 118, 10393.
- [26] M. S. Prévot, X. A. Jeanbourquin, W. S. Bourée, F. F. Abdi, D. Friedrich, R. van de Krol, N. Guijarro, F. Le Formal, K. Sivula, *Chem. Mater.* 2017, 29, 4952.
- [27] B. A. De Angelis, M. Schiavello, J. Solid State Chem. 1977, 21, 67.
- [28] R. Irani, P. Plate, C. Höhn, P. Bogdanoff, M. Wollgarten, K. Höflich, R. van de Krol, F. F. Abdi, J. Mater. Chem. A 2020, 8, 5508.
- [29] Y. Hermans, S. Murcia-López, A. Klein, R. van de Krol, T. Andreu, J. R. Morante, T. Toupance, W. Jaegermann, *Phys. Chem. Chem. Phys.* **2019**, *21*, 5086.
- [30] D. Bae, B. Seger, P. C. K. Vesborg, O. Hansen, I. Chorkendorff, *Chem. Soc. Rev.* 2017, 46, 1933.
- [31] I. A. Digdaya, G. W. P. Adhyaksa, B. J. Trzesniewski, E. C. Garnett, W. A. Smith, *Nat. Commun.* **2017**, *8*, 15968.
- [32] H.-J. Lewerenz, J. Electroanal. Chem. 1993, 365, 121.
- [33] A. J. Bard, A. B. Bocarsly, F. F. Fan, E. G. Walton, M. S. Wrighton, J. Am. Chem. Soc. 1980, 102, 3671.
- [34] S. W. Lee, Y. S. Lee, J. Heo, S. C. Siah, D. Chua, R. E. Brandt, S. B. Kim, J. P. Mailoa, T. Buonassisi, R. G. Gordon, *Adv. Energy Mater.* **2014**, *4*, 1301916.
- [35] J. Lu, J. Aarik, J. Sundqvist, K. Kukli, A. Hårsta, J. O. Carlsson, J. Cryst. Growth 2005, 273, 510.
- [36] X. Ding, X. Wang, Y. Feng, W. Shen, L. Liu, Jpn. J. Appl. Phys. 2020, 59, SGGB16.
- [37] A. Kuzmin, M. Zubkins, R. Kalendarev, Ferroelectrics 2015, 484, 49.
- [38] I. J. McColm, R. Steadman, A. Howe, J. Solid State Chem. 1970, 2, 555.
- [39] Materials Project, https://www.osti.gov/servlets/purl/1322198 (accessed: December 2020).
- [40] W. Smekal, W. S. M. Werner, C. J. Powell, Surf. Interface Anal. 2005, 37, 1059.
- [41] S. Gates-Rector, T. Blanton, Powder Diffr. 2019, 34, 352.
- [42] M. Gorgoi, S. Svensson, F. Schäfers, G. Öhrwall, M. Mertin, P. Bressler, O. Karis, H. Siegbahn, A. Sandall, H. Rensmo, W. Doherty, C. Jung, W. Braun, W. Eberhardt, *Nucl. Instrum. Methods Phys. Res., Sect. A* **2009**, *601*, 48.
- [43] F. Schaefers, M. Mertin, M. Gorgoi, Rev. Sci. Instrum. 2007, 78, 123102.
- [44] G. Kresse, J. Furthmüller, Comput. Mater. Sci. 1996, 6, 15.
- [45] G. Kresse, J. Furthmüller, Phys. Rev. B 1996, 54, 11169.
- [46] G. Kresse, J. Hafner, Phys. Rev. B 1994, 49, 14251.
- [47] J. P. Perdew, K. Burke, M. Ernzerhof, Phys. Rev. Lett. 1996, 77, 3865.
- [48] P. E. Blöchl, Phys. Rev. B 1994, 50, 17953.
- [49] H. J. Monkhorst, J. D. Pack, Phys. Rev. B 1976, 13, 5188.

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- [50] A. Kuzmin, A. Anspoks, A. Kalinko, J. Timoshenko, R. Kalendarev, Sol. Energy Mater. Sol. Cells 2015, 143, 627.
- [51] J. Heyd, G. E. Scuseria, J. Chem. Phys. 2003, 118, 8207.
- [52] M. Harb, L. Cavallo, ACS Omega 2018, 3, 18117.
- [53] S. Lardhi, L. Cavallo, M. Harb, J. Phys. Chem. C 2018, 122, 18204.
- [54] S. Lardhi, A. Curutchet, L. Cavallo, M. Harb, T. Le Bahers, Phys. Chem. Chem. Phys. 2017, 19, 12321.
- [55] D. Noureldine, S. Lardhi, A. Ziani, M. Harb, L. Cavallo, K. Takanabe, J. Mater. Chem. C 2015, 3, 12032.
- [56] M. Harb, D. Masih, S. Ould-Chikh, P. Sautet, J.-M. Basset, K. Takanabe, J. Phys. Chem. C 2013, 117, 17477.
- [57] M. Harb, P. Sautet, E. Nurlaela, P. Raybaud, L. Cavallo, K. Domen, J.-M. Basset, K. Takanabe, Phys. Chem. Chem. Phys. 2014, 16, 20548.
- [58] M. Harb, P. Sautet, P. Raybaud, J. Phys. Chem. C 2011, 115, 19394.